

15295 Regolith Breccia 947 grams

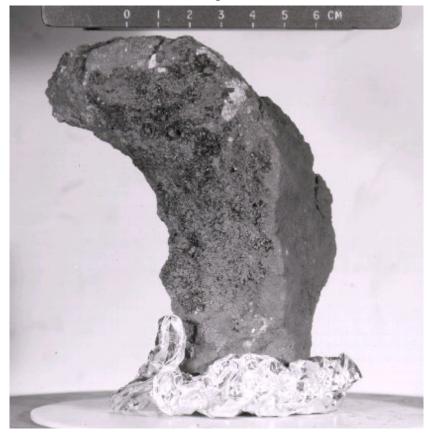


Figure 1: Photo of lunar regolith breccia 15295. NASA photo # S71-51701. Scale in cm.

Introduction

15295 is a brown glass matrix breccia made of local soil components. It was one of several breccias, all similar, collected at station 6, near Spur carter, part way up the slope of Hadley-Delta.

Part of the surface of 15295 is covered with a frothy black glass coating (figure 1). Interior veins of flow-banded green glass were reported by Wilshire and Moore (1974).

15295 is one of the soil breccias studied carefully by Simon et al. (1986) and McKay et al. (1989). The Apollo 15 catalog by Ryder (1985) contains additional information.

Petrography

Simon et al. (1986) compares the mineralogical mode, pyroxene and olivine analyses (figure 5), plagioclase

analyses, etc. for mineral fragments in 15295 with that of other Apollo 15 breccias. They found that the abundance of calcic plagioclase was the best indication of added highland component. Simon et al. also studied the glass compositions of the Apollo 15 breccias, but grouped the data for all the rocks together. In any case, they showed that the regolith breccias from Apollo 15, station 6, contain numerous glasses from various fire fountains (compositional clusters), such as formed the Apollo 15 green glass (Meyer et al. 1975, Delano and Livi 1981).

Warren (1993) tabulates five clasts found in 15295.

Anothosite clast

Warren and Wasson (1978) studied a "large" anorthositic clast (20 x 7 x 9 mm) in 15295 (figures 1 and 2). The texture is described as "cataclastic" with mostly plagioclase ($An_{95.5}$), and sparse, tiny, pyroxenes



Figure 2: Close-up photo of 15295 showing glass coating, brown glass matrix and "large" chalky white clast (anorthosite). NASA S86-39938. Cube is 1 inch, but scale is in cm.

 $(En_{41}Fs_{17}Wo_{42})$. Ir is low, so clast is considered "pristine".

Norite clast ,67

Lindstrom et al. (1989) analyzed a relatively coarse-grained (1.6 mm!) norite clast with plagioclase (An_{94}), pyroxene (En_{76}) and trace troilite, apatite, chromian rutile and a silica mineral reported (figure 6).

Norite clast ,85

Lindstrom et al. (1989) also studied a second norite clast, 30% plagioclase (An_{94}) and 70% pyroxene (En_{75}) (figure 6).

Chemistry

Wanke et al. (1977), Simon et al. (1986), Korotev in McKay et al. (1989) give analysis of matrix of 15295 (table 1, figure 7). Warren and Wasson (1978), Lindstrom et al. (1989) also analyzed several clasts.

Using the composition of 15295, Simon et al. (1986) calculate that it could be a mix of 6% anorthosite, 36% mare basalt, 40% LKFM, 21% KREEP with 0% green glass.

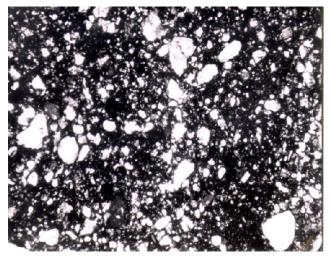


Figure 3: Photomicrograph of thin section of 15295 showing glassy martix. Field of view about 5 mm.

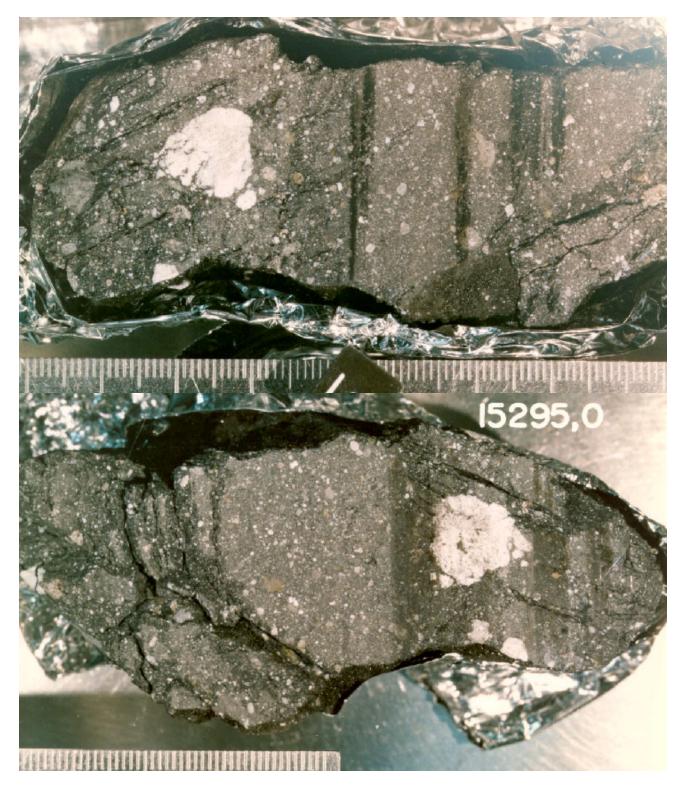


Figure 4: Photos of opposing sawn surfaces of 15295,0 and 15295,46. Scale is 1 cm. NASA S87-43485 and S87-43490. The large white clast is probably one of the norite clasts studied by Lindstrom et al. (1989).

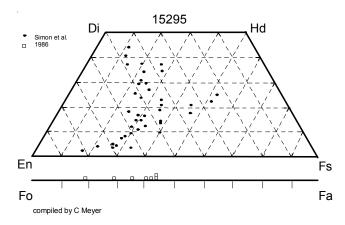


Figure 5: Olivine and pyroxene composition of mineral fragments in matrix of 15295 (replotted from Simon et al. 1986).

gabbronorites gabbronorites ferroananorthosite Lindstrom et al. 1989

90

95

An in plagioclase

75

Figure 6: Plagioclase-pyroxene diagram for pristine rocks from lunar highlands showing norite clasts found in 15295 (Lindstrom et al. 1989).

Other Studies

The rare gas content and isotopic composition of 15295 were determined by Bogard (in McKay et al. 1989).

Processing

15295 was chosen as one of the samples to be studied by the "Regolith Breccia Initiative" (Fruland 1983). It initially broke into several pieces in 1971. A saw cut was made in 1987.

List of Photo #s S71-51701

S86-39938 glass splash and anorthosite clast

S87-43485 sawn surface a norite clast

S87-43490 sawn surface

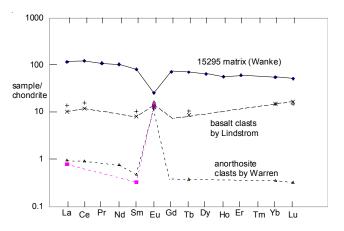


Figure 7: Normalized rare-earth-element diagram for matrix and clasts in 15295 (data from table 1).

Table 1. Chemical composition of 15295.

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reference weight SiO2 % TiO2 Al2O3 FeO MnO MgO CaO Na2O K2O P2O5 S % sum	matrix Wanke 77 ,20 46.68 1.48 16.29 11.87 0.23 10.24 11.33 0.5 0.22 0.22 0.06	matrix McKay 89 ,30 11.5 10.8 0.48	matrix Simon 86 ,35 1.6 16.6 11.8 0.155 11.7 11.1 0.49 0.21	white clast Warren 78 43.9 35.5 0.23 0.18 19.5 0.402	clast Warren 90 45.14 <0.08 34.58 0.35 0.008 0.3 19.48 0.41 0.01	clast Lindstrom ,66 54.3 0.24 11.5 9.8 15.9 7.1 0.26 0.042	clast n 89 ,86 52.1 0.22 12 9.15 17.8 7.2 0.24 0.044
Sc ppm V	24.7	22.2	23.7	0.38	5.3	18.5	15.7
Cr Co Ni Cu Zn Ga Ge ppb As Se Rb	77 2440 39.4 250 4.72 18 4.17 500 0.023 0.15 5.7	2150 38.4 222	85 2360 39 190	17.8 1.4 <15 25.2 3.97 8.2	0.31 0.17 2 3.5 1.9	3830 152 100	3270 18.7 76
Sr Y	135 101	150	120		191		
Zr Nb Mo Ru Rh Pd ppb Ag ppb Cd ppb In ppb Sn ppb Sb ppb Te ppb Cs ppm Ba La Ce Pr Nd Sm Eu Gd Tb Dy Ho Er Tm Yb Lu Hf Ta W ppb Re ppb Os ppb Ir ppb Pt ppb Au ppb Th ppm U ppm technique	394 28	380	340		45		
	0.27 279 27.7 74.3 9.83	0.34 266 24.1 63	270 26 65	0.19	0.054 8.9 0.23 0.56	41 2.44 7.3	55 3.24 9.7
	47 12.1 1.47 14.2	34 11.1 1.41	42 11.9 1.45	0.049 0.78	0.35 0.072 0.94	1.21 0.74	1.54 0.66
	2.58 15.9 3.2 9.75	2.25	2.5 15.7 3.4		0.014	0.32	0.39
	9.07 1.26 9.65 1.17 550	7.8 1.11 9.4 1.08	1.3 8.6 1.22 8.25 1.1		0.058 0.008 1.03 <0.028	2.5 0.42 1.07 <0.13	2.45 0.37 1.6 0.16
	0.71	6.8		0.021	0.022 0.03 0.05		
	2.9 3.89 1.04 (a) INAA	2.2 4 1.08	4.25 1.1	0.041	0.094 0.026	0.34 0.34	0.51 0.2

Lunar Sample Compendium C Meyer 2004